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Щ	UNIVERSITY	ENGINEERING AND COMPUTER	OF COMPUTER
	OF OSTRAVA	SCIENCE	SCIENCE

# Data Visualization

460-4120

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# Smoothed Particle Hydrodynamics (SPH)

- Developed by Gingold, Monaghan and Lucy in 1977 for astrophysical problems
- Used (but not only) for simulation of fluid dynamics
- The fluid is represented by a particle system
- The key idea is to determine some particle properties by taking an average over neighboring particles (similarity with grid less interpolation methods)



Data Visualization

# Grid-based Methods vs SPH

- Traditionally, numerical methods for hydrodynamics are based on discretization (spatial grid, inherently Eulerian)
  - Well developed technique
  - Easy to deal with boundary conditions



- Resolution follows density
- Difficult to implement boundary conditions

Fall 2024





#### Fundamental Idea Behind SPH

• How to compute density from a collection of point masses?







Method 1: construct a mesh around the points, then sum particles within cell and divide by cell volume Method 2: construct local sample volumes, then sum particles within volumes and divide by volume **Method 3**: weight contributions according to distance from sample point (SPH)

#### Particle Description

- Each particle is described by a list of its properties
  - Carried by particle: mass  $m_i$  (const.), position  $x_i$ , velocity  $v_i$ ,
  - Computed: force  $F_i$ , density  $\rho_i$  (varies), pressure  $P_i$ , color  $C_i$
  - These properties forms a particle state vector  $(m_i, x_i, v_i, F_i, \rho_i, P_i, C_i)$

# Density From Collection of Masses

• Density  $\rho_i$  at the point  $x_i$  is computed (approximated) as a weighted sum of N particle masses  $m_j$  as follows

$$\rho_i \approx \sum_{j=1}^N m_j W_{ij}$$

where the smoothing kernel

$$W_{ij} = W(\|\boldsymbol{x}_i - \boldsymbol{x}_j\|, h)$$

and h is smoothing length

• This formulation ensures that the resolution follows density such that  $\rho h^3 = const.$ 

#### Smoothed Particle Interpolation

- If we know that  $\rho_i = \sum_{j \in N_i} m_j W_{ij}$ , how we can compute arbitrary (smoothed) physical quantity  $A_i$  of *i*-th particle?
- We can use a volume of *i*-th particle  $V_i = \frac{m_i}{\rho_i}$  from which  $m_i = V_i \rho_i$
- Now we get  $\rho_i = \sum_{j \in N_i} V_j \rho_j W_{ij}$  and we have the same quantity on both sides of equation provided that  $i \in N_i$

• It also holds that 
$$V_i = \frac{m_i}{\rho_i} = \frac{m_i}{\sum_{j \in N_i} m_j W_{ij}}$$

• Now, in general, for arbitrary quantity  $A_i$ , we can write the smoothed interpolation as follows

$$A_{i} = \sum_{j \in N_{i}} V_{j} A_{j} W_{ij} = \sum_{j \in N_{i}} \frac{m_{j}}{\sum_{k \in N_{j}} m_{k} W_{jk}} A_{j} W_{ij}$$

### Kernel Function

- Kernel function should approximate a delta function, i.e. particles which are closer should contribute more to the local evaluation of fluid properties
- First choice Gaussian kernel

$$W(r,h) = \frac{1}{\pi^{1/3}h^3} e^{-(r/h)^2}$$

- One issue with this function is that the support is not compact and summation must be done over all particles
- One can choose a kernel with compact support (weight vanishes beyond a given distance)
- Better choice Cubic spline kernel (neighborhood contains only particles lying within 2h distance)

### Kernel Function

- In general, smoothing kernel can be any function which satisfies the following to properties
  - Normalization

$$\int_{V} W(r,h) \mathrm{d}V = 1$$

• Dirac delta function approximation

$$\lim_{h\to 0} W(r,h) = \delta(r)$$

# Cubic Spline Kernel Function

• The smoothing kernel function is defined as follows

$$W_{ij}(r,h) = \frac{3}{2\pi h^3} \begin{cases} \frac{2}{3} - q^2 + \frac{1}{2}q^3 & 0 \le q < 1\\ \frac{1}{6}(2-q)^3 & 1 \le q < 2\\ 0 & q \ge 2 \end{cases}$$
  
where  $q = \frac{r}{h} = \frac{\|x_i - x_j\|}{h}$ 

#### Arbitrary Quantity Derivatives

- Repeat that  $A_i = \sum_{j \in N_i} V_j A_j W_{ij}$
- Gradient

$$\nabla A_i = \sum_{j \in N_i} V_j A_j \nabla_i W_{ij}$$

• Laplacian

$$\Delta A_i = \sum_{j \in N_i} V_j A_j \Delta_i W_{ij}$$

#### Kernel Gradient

• 
$$\nabla_i W_{ij} = \begin{bmatrix} \frac{\partial W_{ij}}{\partial x_i} \\ \frac{\partial W_{ij}}{\partial y_i} \\ \frac{\partial W_{ij}}{\partial z_i} \end{bmatrix} = \frac{\partial W_{ij}}{\partial q} \nabla_i q = \frac{\partial W_{ij}}{\partial q} \frac{x_i - x_j}{\|x_i - x_j\| h}$$

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Note that 
$$\frac{\partial}{\partial x} \left( \frac{|x|}{h} \right) = \frac{x}{|x|h} = \frac{x}{\sqrt{x^2}h} = \frac{\operatorname{sgn} x}{h}$$

$$\bullet \frac{\partial W_{ij}}{\partial q} = \frac{3}{2\pi h^3} \begin{cases} -2q + \frac{3}{2}q^2 & 0 \le q < 1\\ -\frac{1}{2}(2-q)^2 & 1 \le q < 2\\ 0 & q \ge 2 \end{cases}$$

# Kernel Laplacian

• 
$$\Delta_i W_{ij} = \frac{\partial^2 W_{ij}}{\partial x_i^2} + \frac{\partial^2 W_{ij}}{\partial y_i^2} + \frac{\partial^2 W_{ij}}{\partial z_i^2} = \frac{\partial^2 W_{ij}}{\partial q^2} \frac{1}{h^2} + \frac{\partial W_{ij}}{\partial q} \frac{2}{h}$$

• 
$$\frac{\partial^2 W_{ij}}{\partial q^2} = \frac{3}{2\pi h^3} \begin{cases} -2+3q & 0 \le q < 1\\ 2-q & 1 \le q < 2\\ 0 & q \ge 2 \end{cases}$$

# Fluid Dynamics

• Govern by Navier-Stokes equation (momentum equation)

$$\boldsymbol{a} = \frac{\partial \boldsymbol{v}}{\partial t} = -\frac{1}{\rho} \nabla p + \frac{\mu}{\rho} \Delta \boldsymbol{v} + \frac{1}{\rho} \mathbf{F}_{ext} + \boldsymbol{g}$$
$$\overset{\boldsymbol{a}^{P}}{\boldsymbol{a}^{P}} = \mathbf{a}^{V} \qquad \mathbf{a}^{E} = \mathbf{a}^{G}$$

- It stems from motion equation in elasticity
- $\boldsymbol{v}$  is a velocity field, p is a pressure field,  $\rho$  is a density field
- Note: dynamic (absolute) viscosity  $\mu$  and kinematic viscosity (momentum diffusivity)  $\nu = \frac{\mu}{\rho}$

# Fluid Dynamics

- Velocities and positions of particles are calculated from acting forces
- Three forces are applied on each particle (external force is excluded here)
  - Gravity force  $F_i^G = m_i g$
  - Fluid pressure force  $\mathbf{F}_i^P = -V_i \sum_{j \in N_i} V_j p_j \nabla_i W_{ij}$
  - Fluid viscosity force  $F_i^V = m_i v \sum_{j \in N_i} V_j v_j \Delta_i W_{ij}$

# Gravity Force

• The steady gain in speed of mass particle caused exclusively by the force of gravitational attraction is give by

0

$$\boldsymbol{F}_{i}^{G} = m_{i}\boldsymbol{g}$$



#### Fluid Pressure Force

• We compute fluid pressure force induced by pressure term from NSE as follows

$$\boldsymbol{F}_{i}^{P} = m_{i}\boldsymbol{a}_{i}^{P} = m_{i}\left(-\frac{1}{\rho_{i}}\nabla p_{i}\right) = -V_{i}\nabla p_{i} = -V_{i}\sum_{j\in N_{i}}V_{j}p_{j}\nabla_{i}W_{ij}$$

$$\underbrace{p_{i}+p_{j}}_{2} \quad \text{action} \neq \text{reaction}$$
symetrization needed

- Pressure force depends on the difference (i.e. gradient) of pressure
- There is no pressure force (i.e. no acceleration) in areas with constant pressure

$$\mathbf{F}_{i}^{P} = \mathbf{0} \qquad \qquad \left\| \mathbf{F}_{i}^{P} \right\| >$$

#### Fluid Pressure Force

• Pressure at  $x_i$  can be computed via the ideal gas state equation

$$p_i = k(\rho_i - \rho)$$

where k is a gas constant that depends on the temperature and  $\rho$  is a rest (reference) density

• Here, we use Cole equation

$$p_i = B\left(\left(\frac{\rho_i}{\rho}\right)^{\gamma} - 1\right)$$

where B is tuneable gas constant with presure units and adiabatic index  $\gamma \cong 7$ 

### Fluid Viscosity Force

• We compute fluid viscosity force induced by viscosity term from NSE as follows

$$F_{i}^{V} = m_{i} \boldsymbol{a}_{i}^{V} = m_{i} \frac{\mu}{\rho_{i}} \Delta \boldsymbol{v}_{i} = m_{i} \nu \Delta \boldsymbol{v}_{i} = m_{i} \nu \sum_{j \in N_{i}} V_{j} \boldsymbol{v}_{j} \Delta_{i} W_{ij}$$
Material density is constant in case of incompressible flow (i.e. resist volume change) but it is not absolutely true here
$$V_{j} \boldsymbol{v}_{j} - \boldsymbol{v}_{i} \text{ asymmetric again force depends on velocity differences only}$$

- Viscosity causes loss of energy due to internal friction
- In viscous flow, particles should move together with the same velocity
- Resulting force is minimizing velocity difference between neighboring particles

#### Particles Position Update

- For every *i*-th particle compute force  $F_i = F_i^P + F_i^V + F_i^G$  using its neighborhood set  $N_i$
- Update the velocity  $v_i += \frac{F_i}{m_i} \Delta t$
- Update the position  $x_i += v_i \Delta t$

#### Kernel Variants

• Other various kernels were developed to improve numerical stability

#### Kernel Variants

• Other various kernels were developed to improve numerical stability

• 
$$\Delta W_{ij}(r,h) = \begin{cases} \frac{45}{\pi h^6}(h-r) & 0 \le r < h \\ 0 & r \ge h \end{cases}$$

• These kernels can be used in the following simplified formulas for computing accelerations from the original NSE

• 
$$\boldsymbol{a}_{i}^{G} = \boldsymbol{g}$$
  
•  $\boldsymbol{a}_{i}^{P} = -\sum_{j \in N_{i}, j \neq i} \left( \frac{p_{i}}{\rho_{i}^{2}} + \frac{p_{j}}{\rho_{j}^{2}} \right) m_{j} \nabla W_{ij}$   
•  $\boldsymbol{a}_{i}^{V} = \frac{\mu}{\rho_{i}} \sum_{j \in N_{i}, j \neq i} (\boldsymbol{v}_{j} - \boldsymbol{v}_{i}) \frac{m_{j}}{\rho_{j}} \Delta W_{ij}$ 

# Algorithm

```
Init particles (m_i = m/n, v_i = 0, x_i \in \text{cube})

For each time step:

Init neighborhoods

Compute densities

Compute pressures from densities

For each i-th particle:

Compute acceleration a_i = a_i^G + a_i^P + a_i^V

Update velocity v_i += a_i \Delta t

Update position x_i += v_i \Delta t

Check boundaries (prevent particle from leaving simulation domain)
```

#### kNN Radius Search

typedef std::vector<int> Neighbourhood;

Neighborhood SPHSolver::GetNeighborhood( const int i, const float r ) {
 Neighborhood indices;
 std::vector<float> distances;

```
Vector3f x_i = particles_[i].position;
std::vector<float> query{ x_i.x, x_i.y, x_i.z };
cvflann::SearchParams search_params;
```

int n = search\_index\_->radiusSearch(query,indices,distances,r\*r,search\_params);

```
if ( n > 0 ) return Neighborhood( &indices[0], &indices[n] );
return Neighborhood( 0 );
```

#### Search Index

typedef cv::flann::GenericIndex<cv::flann::L2<float>> SearchIndex;

```
int SPHSolver::InitSearchIndex() {
```

```
cv::Mat features( n_, 3, CV_32FC1 );
```

```
for ( int i = 0; i < n_; ++i ) {</pre>
```

```
const Vector3f & position = particles_[i].position;
```

```
features.at<float>( i, 0 ) = position.x;
features.at<float>( i, 1 ) = position.y;
features.at<float>( i, 2 ) = position.z;
```

```
}
search_index_ = std::make_unique<SearchIndex>( features,
```

```
cvflann::KDTreeSingleIndexParams( 10, false ) );
```

### Parameters Settings

- $n = 40 \times 40 \times 40$  particles
- $\rho = 1000 \text{ kg} \cdot \text{m}^{-3}$  (rest density)
- $g = (0,0,-9.81) \text{ m} \cdot \text{s}^{-2}$
- B = 3.0 (fluid stiffness)
- h = 0.055 m (smoothing length)
- Simulation domain size  $1\times1\times1$  m
- Initial particle separation 0.018 m
- $\mu = 3.5 \text{ N} \cdot \text{s} \cdot \text{m}^{-2}$  (dynamic or absolute viscosity)
- $\Delta t = 0.001 \text{ s}$

#### SPH Results



#### SPH Results



#### Iso Surface Reconstruction



- Originally published in Marching cubes: A high resolution 3D surface construction algorithm". ACM SIGGRAPH Computer Graphics (1987)
- Topological issues fixes and further improvements were presented later
- The goal is to extract a polygonal mesh of an isosurface from the 3D discrete scalar field
- Elements of such 3D scalar field are called voxels (CT and MRI scans)
- In general, the algorithm determines the polygons needed to represent the part of isosurface passing by the give voxel (cube)

- 1. Assign a scalar value to each vertex of a cube
- If the scalar field value at the given vertex is bellow a certain threshold (isovalue), assign 0 to the appropriate bit, otherwise set this bit to 1



In total, we get 2<sup>8</sup> = 256 possible assignments (two states {inside, outside} in 8 vertices of a cube)

- 2. Based on step 1, we connect points on 12 line segments of the cube
- Some lookup table contains 256 entries of 12 bits representing connected midpoints on line segments



 Of the 256 different combinations 2 will not give any object, 8 will result in a triangle placed in every corner of the cube and so on

- 3. Set proper positions of all vertices  $x_{iso}$  on segments  $x_i$ ,  $x_j$  via interpolation
- Interpolation weights are derived from known functional values at selected vertices of given cube

$$\Delta f = \frac{f_{iso} - f_i}{f_j - f_i}$$

• Final point for each vertex of the triangle is computed as follows

$$\boldsymbol{x}_{iso} = \boldsymbol{x}_i + (\boldsymbol{x}_j - \boldsymbol{x}_i) \Delta f$$

Used indexing: Note that vertex and edge indices may differ in a particular implementation, e.g. https://paulbourke.net/geometry/polygonise/



#### Marching Cubes Results



**Original SPH particles** 

Marching Cubes iso-surface reconstructions with two different sizes of sampling steps producing different amount of details

Note that presented images are captured at different times

# References and Further Readings

- MÜLLER, Matthias; CHARYPAR, David; GROSS, Markus H. Particle-based fluid simulation for interactive applications. In: *Symposium on Computer animation*. 2003.
- IHMSEN, Markus, et al. SPH fluids in computer graphics. 2014.
- KOSCHIER, Dan, et al. Smoothed particle hydrodynamics techniques for the physics based simulation of fluids and solids. 2024.
- VIOLEAU, Damien; ISSA, Reza. Numerical modelling of complex turbulent free-surface flows with the SPH method: an overview. *International Journal for Numerical Methods in Fluids*, 2007.
- SUTTI, Marco. SPH treatment of boundaries and application to moving objects. École polytechnique fédérale de Lausanne. 2014.

#### **Technical Notes**

Expected directory structure:



Created VS solution

• How to set paths for OpenCV and VTK libraries...

MC Property Pages		? ×	MC Property Pages		? ×	
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Optimization	All Header Files are Public	No		Amp Default Accelerator	WARP software accelerator	
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# **Technical Notes**

#### • How to set paths for OpenCV and VTK libraries...



#### **Technical Notes**

